

# Gravitation and the $1/r^2$ law: why distance is squared

Version PDF de travail - contenu du cours HTML canonique

## Course compass

Guiding question: Why does gravity decrease quickly with distance?

## Essential answer

Gravitation and the  $1/r^2$  law: why distance is squared. The question to solve is: Why does gravity decrease quickly with distance? We start from the concrete problem before notation. The goal is to understand what we seek, then why mathematics becomes useful. The rest of the course turns that idea into an auditable line of reasoning: explicit units, stated assumptions, reproducible calculations, order-of-magnitude checks and interpretation limits. A result is useful only when the reader can explain what it measures, where every input came from and which engineering decision it can support.

Starting synthesis: derivations, examples, limitations and sources are developed in the course body.

## Key terms before you start

acceleration · force · mass · gravity · unit

## 1 — The physical question

Why does gravity decrease quickly with distance?

We start from the concrete problem before notation. The goal is to understand what we seek, then why mathematics becomes useful.

## 2 — How to read the symbols and units

$F = G \times m_1 \times m_2 / r^2$ . Read: "F equals G times m one times m two, divided by r squared".

F — gravitational force ; G — universal gravitational constant ;  $m_1, m_2$  — masses ; r — distance between centres

## 3 — Where does the relation come from?

In Newton's model, attraction between two masses grows with each mass and decreases with the square of the distance between their centres. The  $1/r^2$  dependence underlies much classical orbital mechanics.

Every number used below is explicitly treated as data, convention, learning assumption, or calculated result.

## 4 — A — Double distance

Where do the numbers come from? Force F at distance r. What at 2r?

Step-by-step calculation:  $F_2/F_1 = r^2/(2r)^2 = 1/4$ .

Doubling distance divides force by four.

## 5 — B — Triple distance

Where do the numbers come from? Same masses, distance 3r.

Step-by-step calculation:  $F_2/F_1 = 1/9$ .

The square makes the decrease rapid.

## 6 — C — Double mass

Where do the numbers come from?  $m_2$  becomes  $2m_2$  at same distance.

Step-by-step calculation:  $F_2/F_1 = 2$ .

Mass is linear in the equation, unlike distance.

## 7 — Sensitivity, inverse calculation, and sanity check

Change one input, predict the direction of the result, calculate, then check units, sign, order of magnitude, and limits.

Essential limit for Gravitation and the  $1/r^2$  law: why distance is squared: the displayed relation is a learning model. A real system adds detailed geometry, variable properties, sensors, uncertainty, transients, and testing.

## 8 — Why this matters in a mission

In a space mission, why does gravity decrease quickly with distance? The useful skill is not reciting the formula but knowing which data are needed, which are measured, and when the model becomes insufficient.

## 10 — Go deeper: from calculation to physical understanding

### Why gravity decreases with the square of distance

In Newton's model, the same gravitational influence spreads geometrically as distance increases. The  $1/r^2$  law means that doubling the centre-to-centre distance divides force by four, while tripling it divides force by nine. The reader should understand this sensitivity before memorising  $G$  or any planetary mass: distance matters strongly.

### Distance $r$ is measured between centres

For approximately spherical bodies,  $r$  is the distance between their centres of mass, not altitude above the ground. At Earth's surface,  $r$  is about Earth's radius; at 400 km altitude those 400 km must be added to the radius. Confusing altitude with centre distance creates a systematic calculation error.

### Why mass appears twice

Force increases with either mass: doubling the first mass doubles  $F$ , and doubling the second does the same. Yet the free-fall acceleration of a small object does not depend on its own mass in the ideal model because we then divide force by that mass through  $F=ma$ . Connecting the two equations removes the apparent paradox.

### From force calculation to orbit calculation

For a space trajectory, engineers mainly use gravitational acceleration and equations of motion, often with several bodies and more precise models. Newton's law remains the intuitive foundation: it explains why Earth dominates near Earth and why the Sun becomes central during interplanetary flight.

## 9 — Exercises and answers

### Challenge 1

Force  $F$  at distance  $r$ . What at  $2r$ ?

### Challenge 2

Same masses, distance  $3r$ .

### Challenge 3

$m_2$  becomes  $2m_2$  at same distance.

## Primary and technical sources

- NASA Basics of Space Flight — Gravity & Mechanics
- NASA Glenn — Newton's Laws of Motion